

HYDRATED MINERALS IN METEORITES

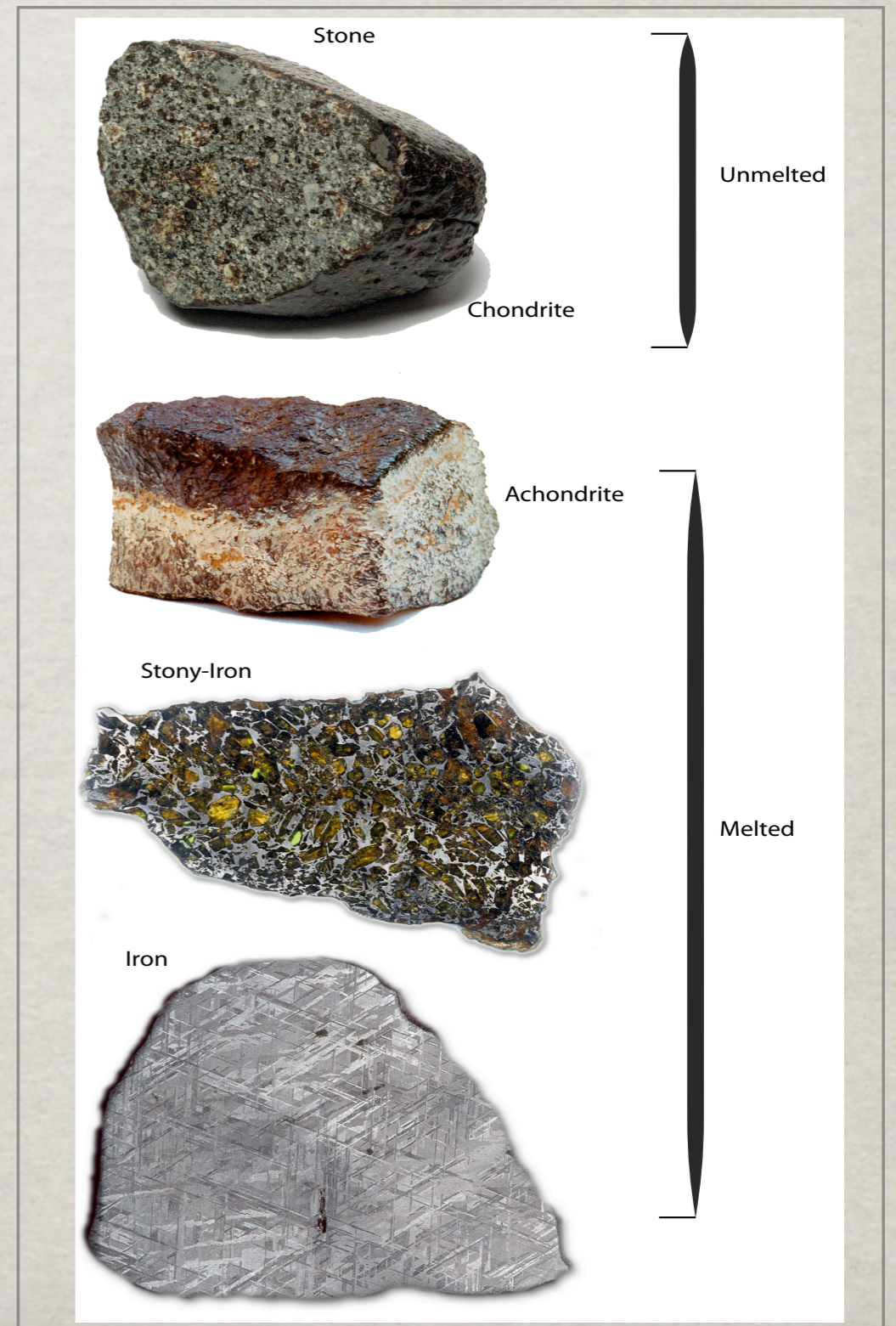
G . K . B E N E D I X , K . T . H O W A R D ,
K . A . D Y L , P . A . B L A N D

OVERVIEW

- ✿ Meteorite Basics
- ✿ Evidence for aqueous alteration
 - ✿ Diagnostic minerals
 - ✿ Distribution in meteorites
 - ✿ Abundances (i.e. modal mineralogy)
- ✿ Implications for spectra

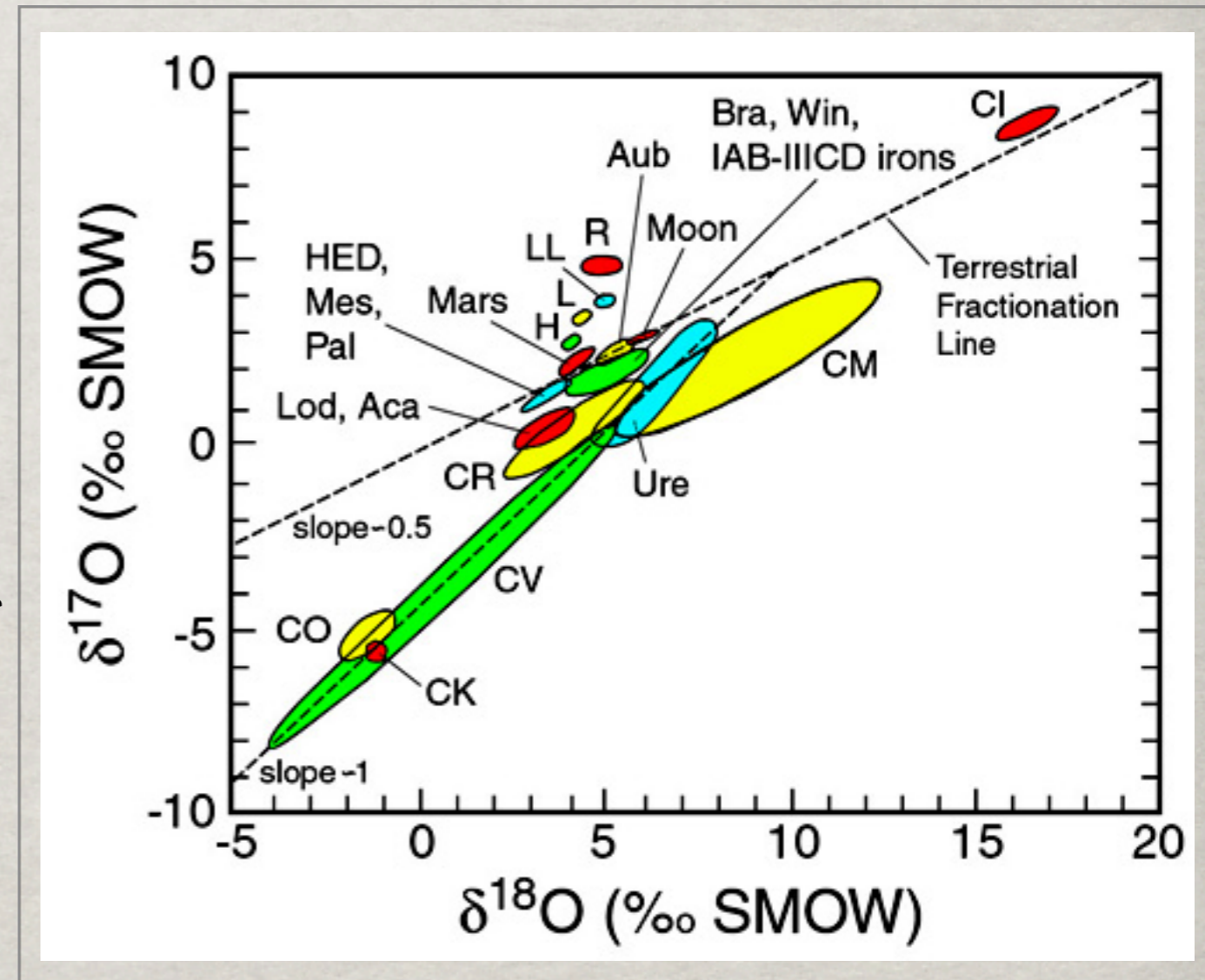
METEORITE CLASSES

- ☀ Three broad categories
- ☀ Range from primitive to differentiated
- ☀ Can be grouped according to chemistry
- ☀ into ~40 classes



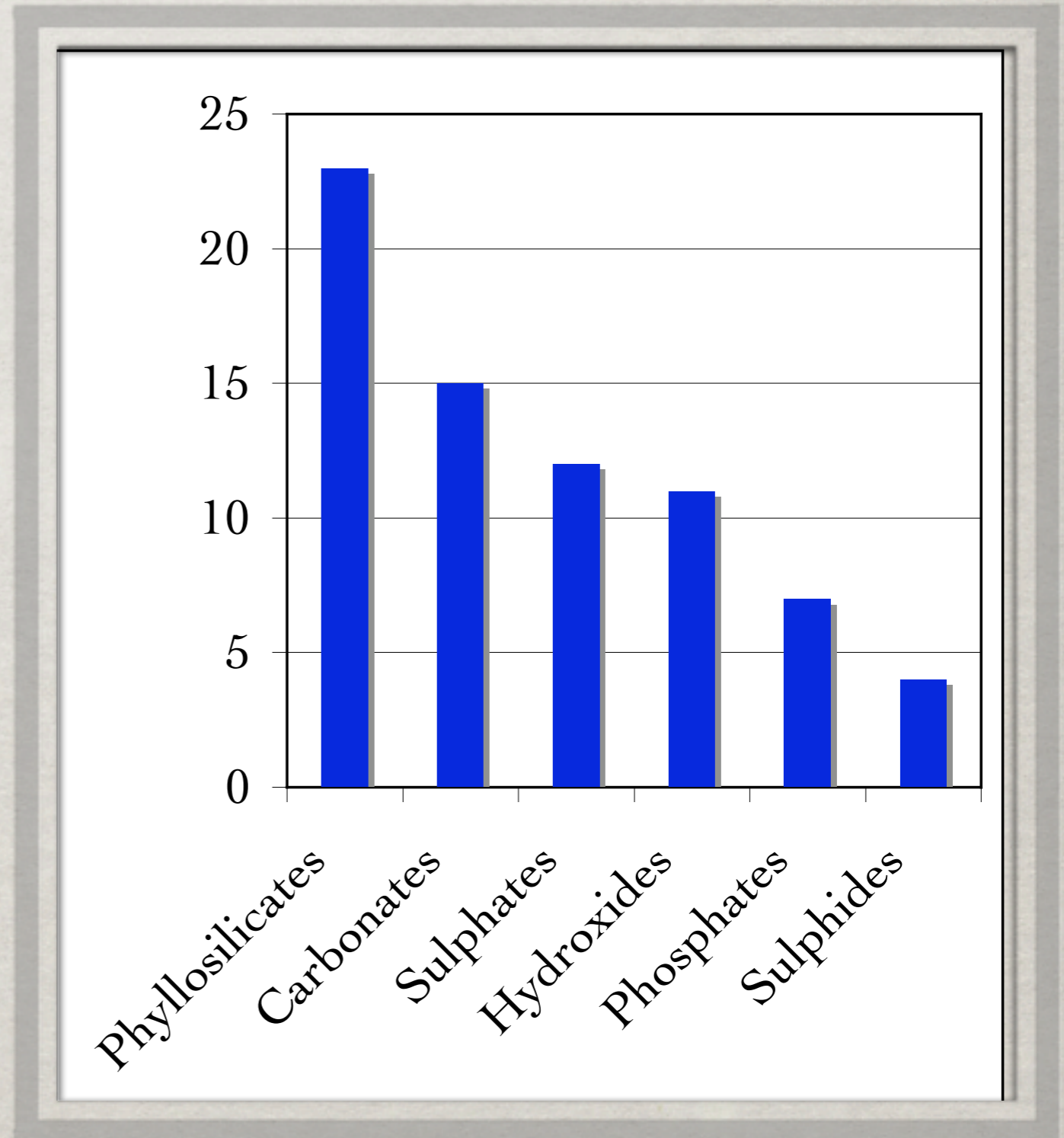
METEORITE CLASSES

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HYDRATED MINERALS

- ☼ 275 different mineral species reported in meteorites (Rubin, 1997)
- ☼ ~78 different hydrated minerals in meteorites



AQUEOUS ALTERATION

- ✱ Most obvious effect of interaction with water is the presence of Fe, Mg phyllosilicates
- ✱ Typical phyllosilicates (ideal formula)
 - ✱ Cronstedtite - $\text{Fe}_2^{+2}\text{Fe}^{+3}(\text{SiFe}^{+3})\text{O}_5(\text{OH})_4$
 - ✱ Chrysotile/Lizardite/Antigorite - $\text{Mg}_3\text{Si}_2\text{O}_5(\text{OH})_4$
 - ✱ Saponite -
 $(\text{Ca,Na})_{0.3}(\text{Mg,Fe}^{+2})_3(\text{Si,Al})_4\text{O}_{10}(\text{OH})_2 \cdot 4\text{H}_2\text{O}$

GENERAL REACTION SEQUENCE

- ✻ H₂O reacts with Fe metal and sulphide to form Tochilinite
(2[(Fe,Mg,Cr,Ni□)S•1.57-1.85[(Mg,Fe,Ni,Al,Ca)(OH)₂])
- ✻ Solution reacts with silicates to form Fe-rich phyllosilicates
- ✻ Mg-phyllosilicates are last to form as solution evolves further



olivine

serpentine

100µm*



Mag = 398 X

WD = 15 mm

EHT = 20.00 kV

Spot Size = 500

Signal A = QBSD

File Name = KH0006.tif

Chamber =

14 Pa

DISTRIBUTION OF MINERALS - CV



A slice through the CV3 meteorite Vigarano - Photo courtesy of NHM

DISTRIBUTION OF MINERALS - CM



A Murchison meteorite specimen at the National Museum of Natural History (Washington)

DISTRIBUTION OF MINERALS - CI

Ivuna, CI Carbonaceous Chondrite



© Natural History Museum, London.

(Image courtesy of the Natural History Museum, London.)

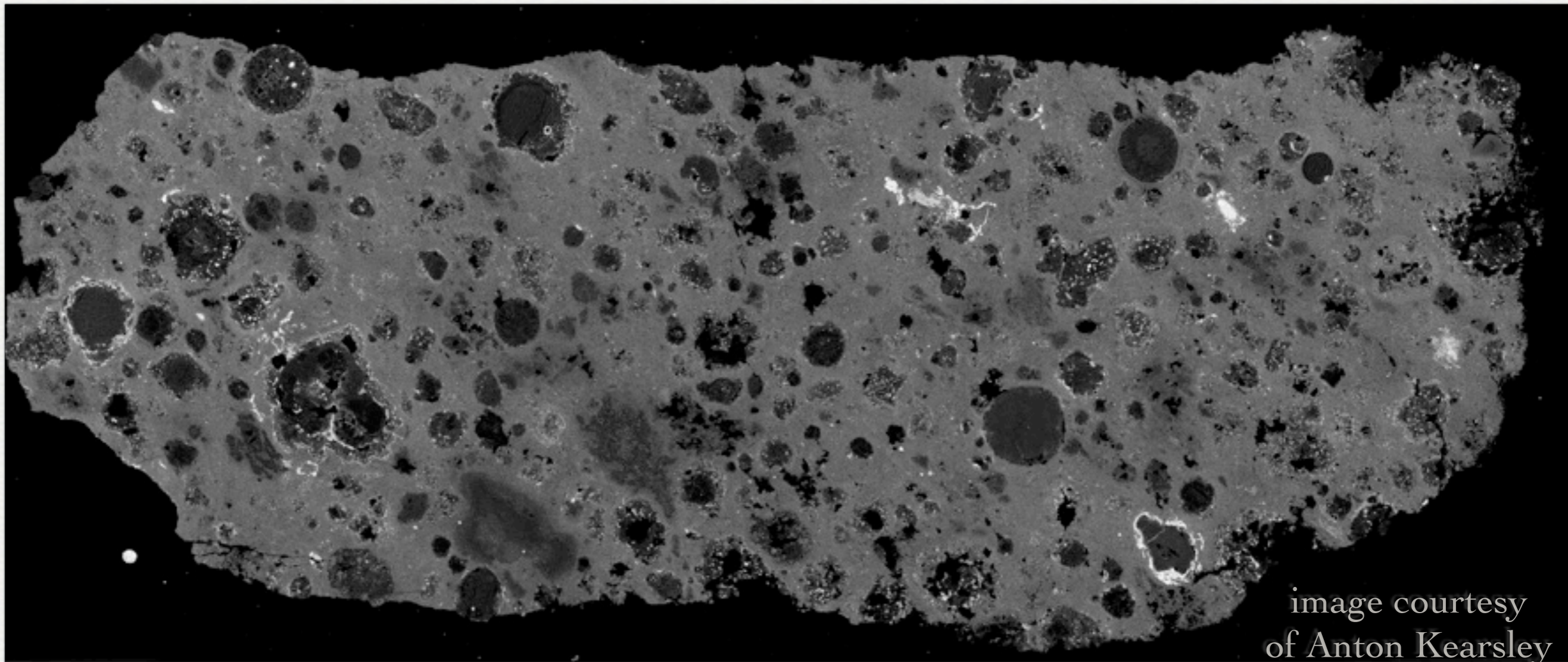
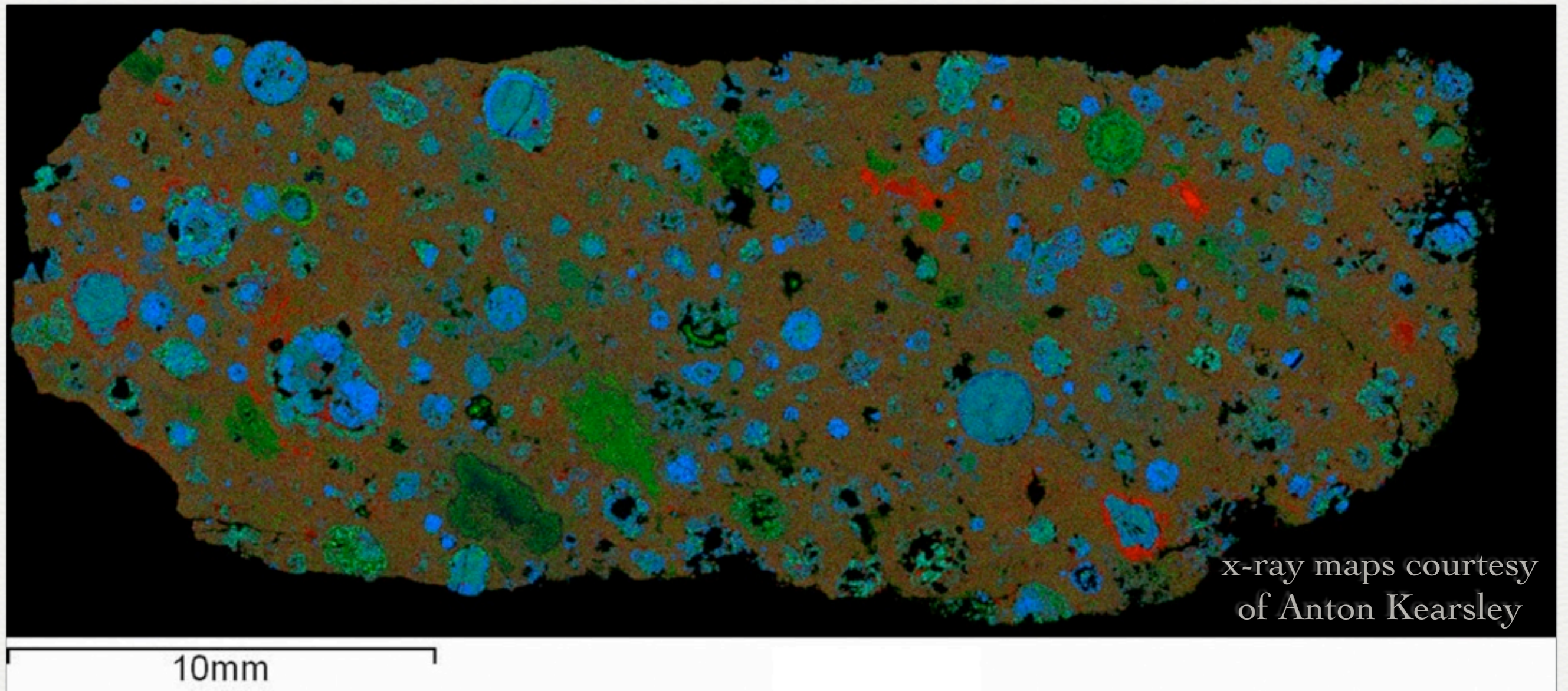


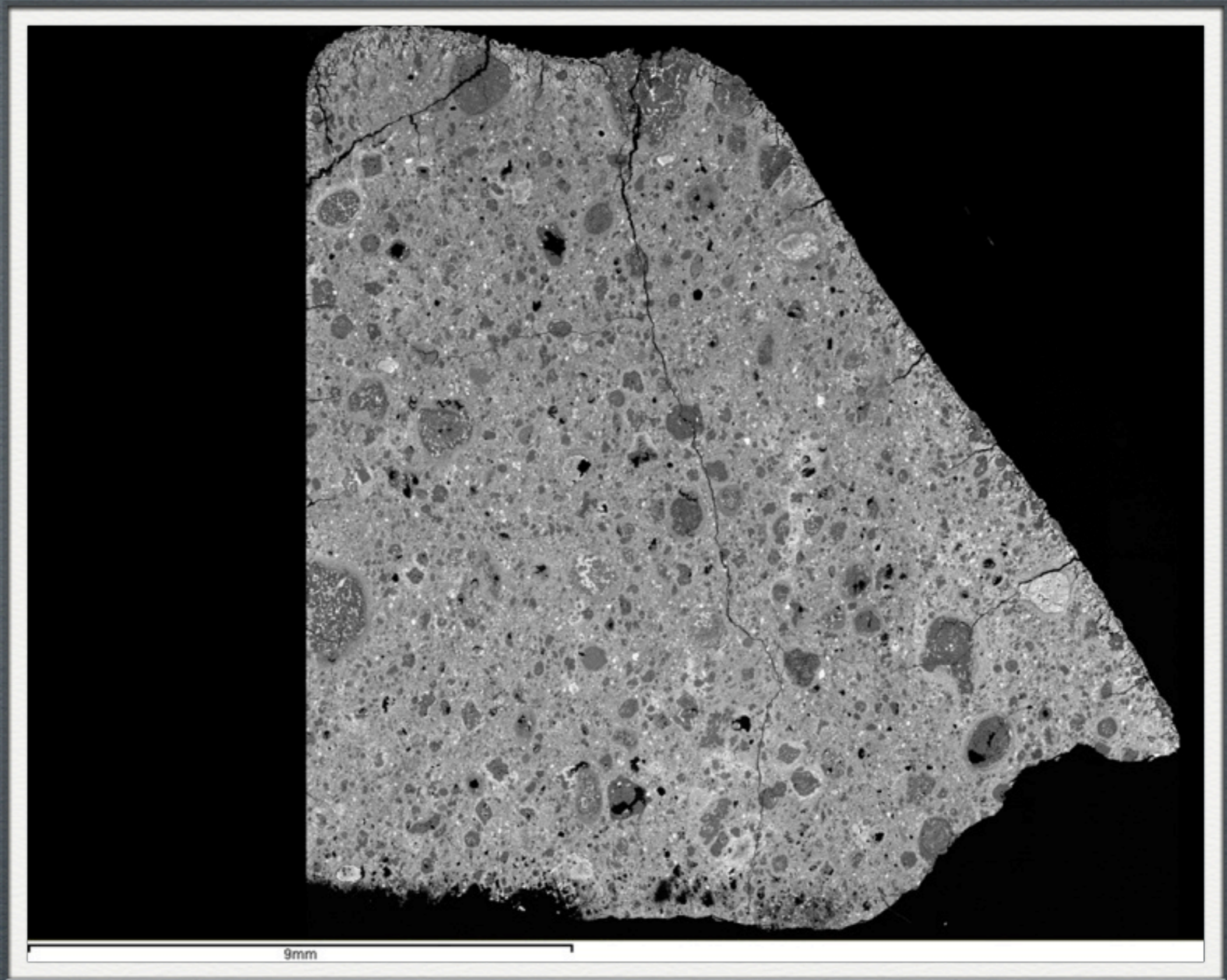
image courtesy
of Anton Kearsley

ALLENDE - CV CHONDRITE

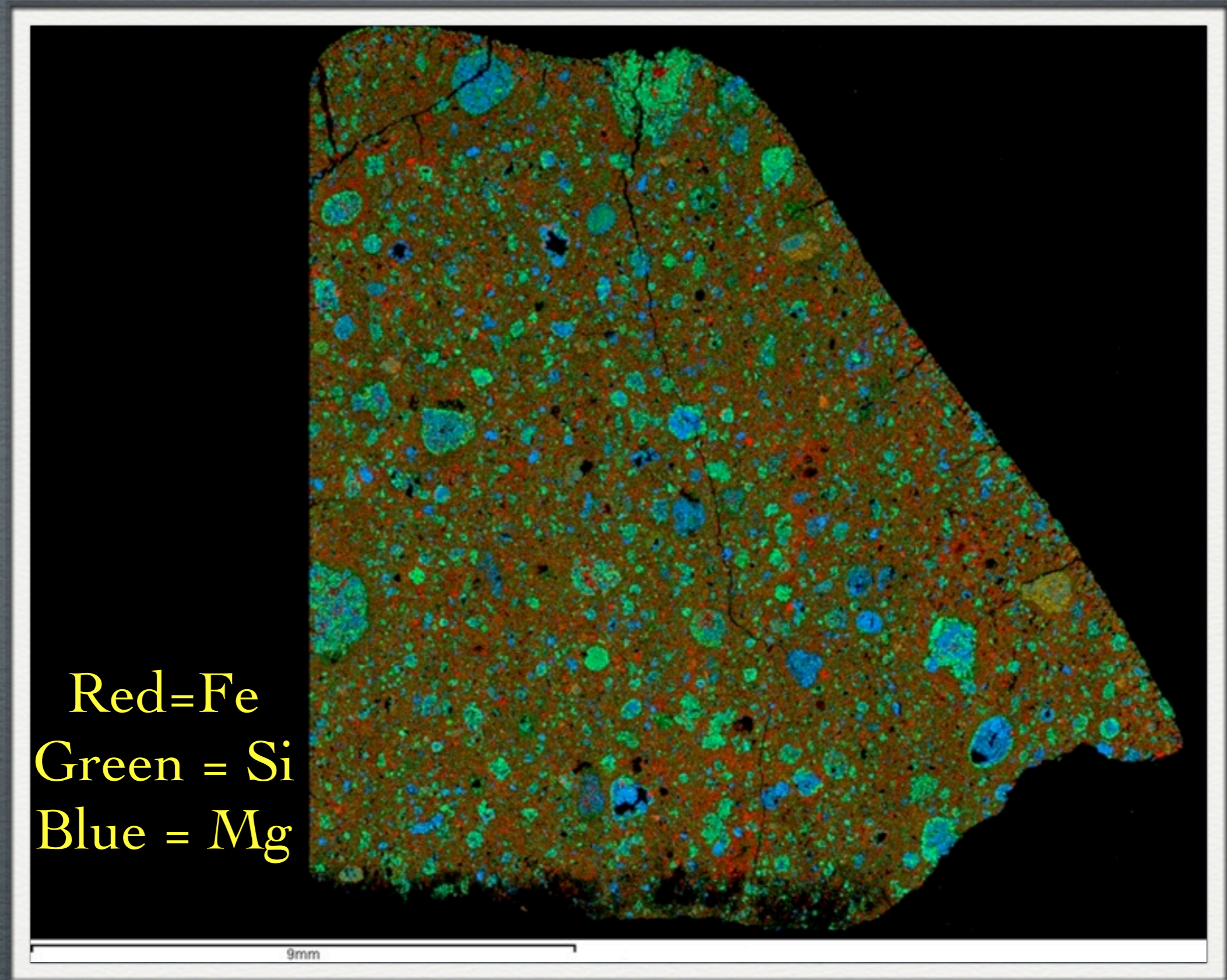


Red=Fe; Green = Si; Blue = Mg

ALLENDE - CV CHONDRITE

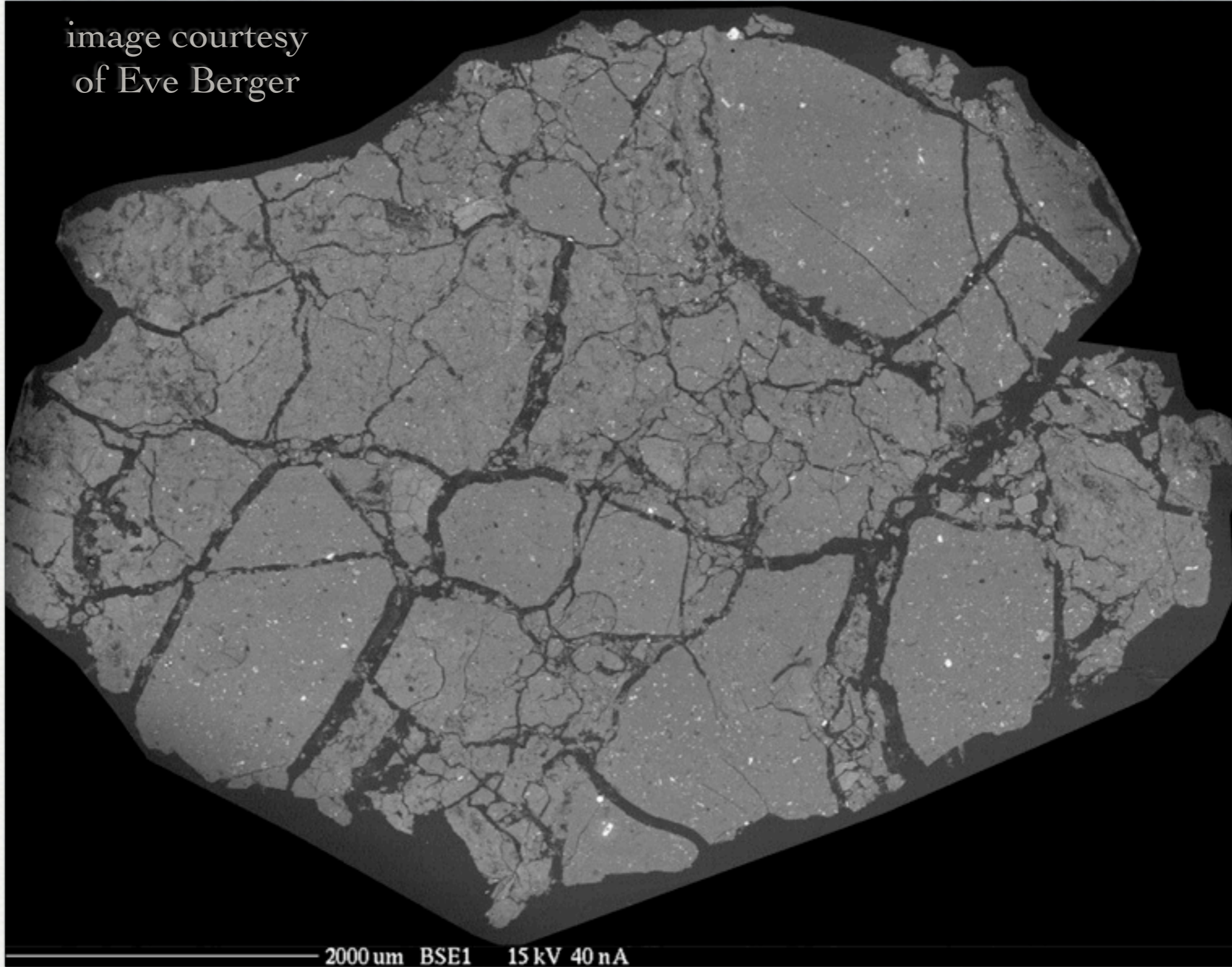


MURCHISON - CM CHONDRITE



MURCHISON - CM CHONDRITE

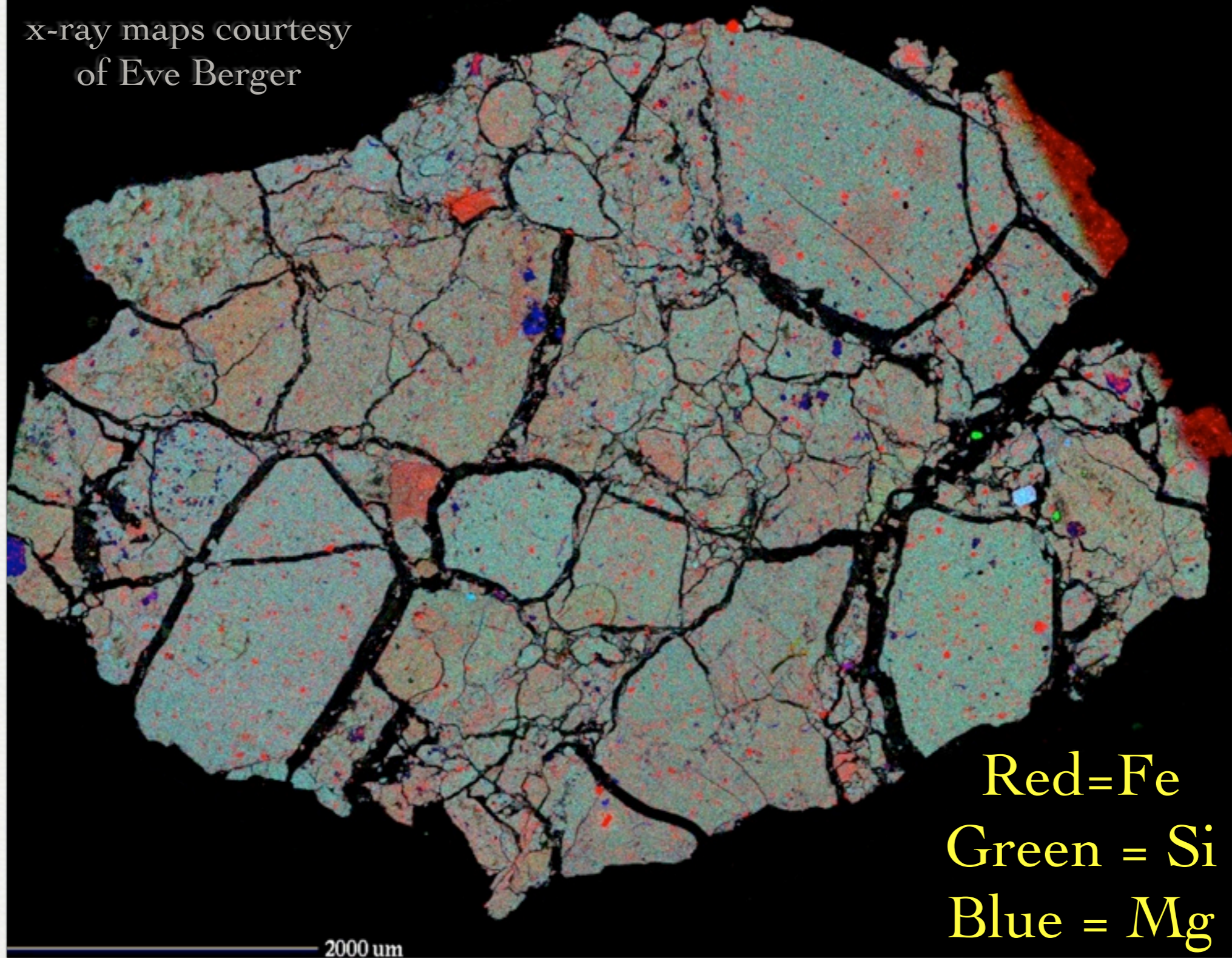
image courtesy
of Eve Berger



2000 um BSE1 15 kV 40 nA

ORGUEIL - CI CHONDRITE

x-ray maps courtesy
of Eve Berger

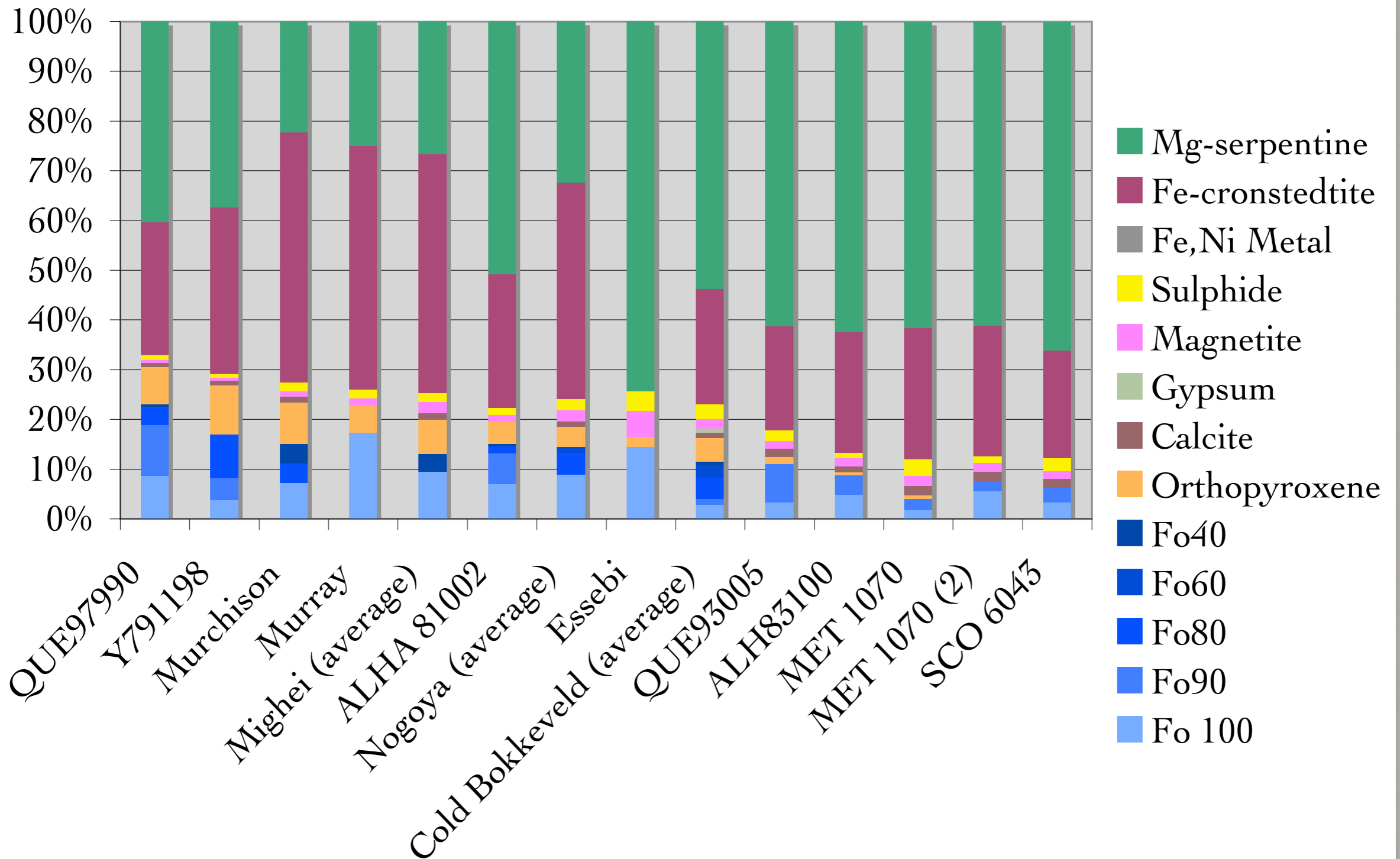


ORGUEIL - CI CHONDRITE

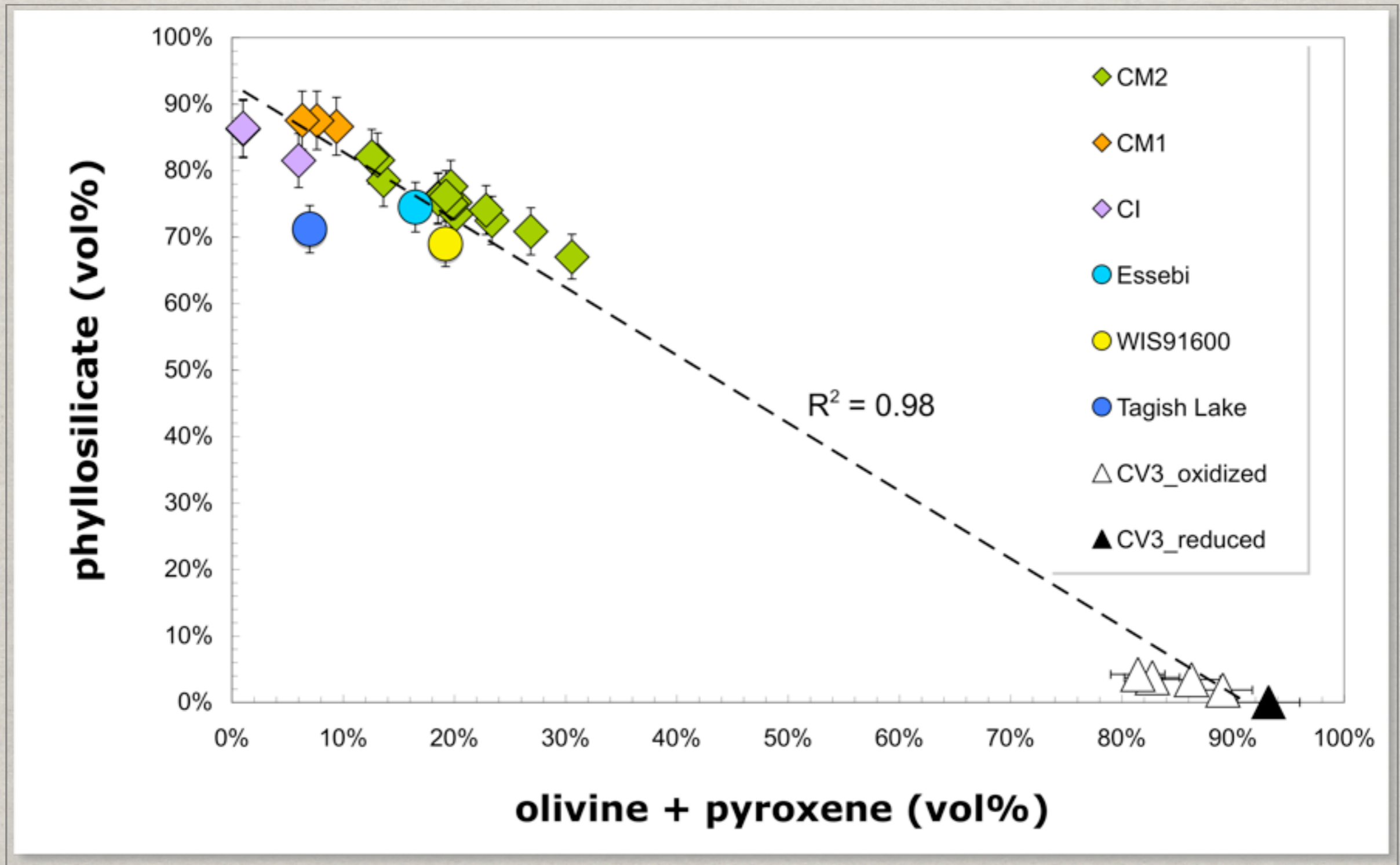
MODAL ABUNDANCE

- ✱ How much is really there?
 - ✱ Can see variation of hydrated minerals in different meteorite types
- ✱ X-Ray Diffraction allows quantification of modal abundances of minerals at $>1\text{vol}\%$

Modal Abundance - CM



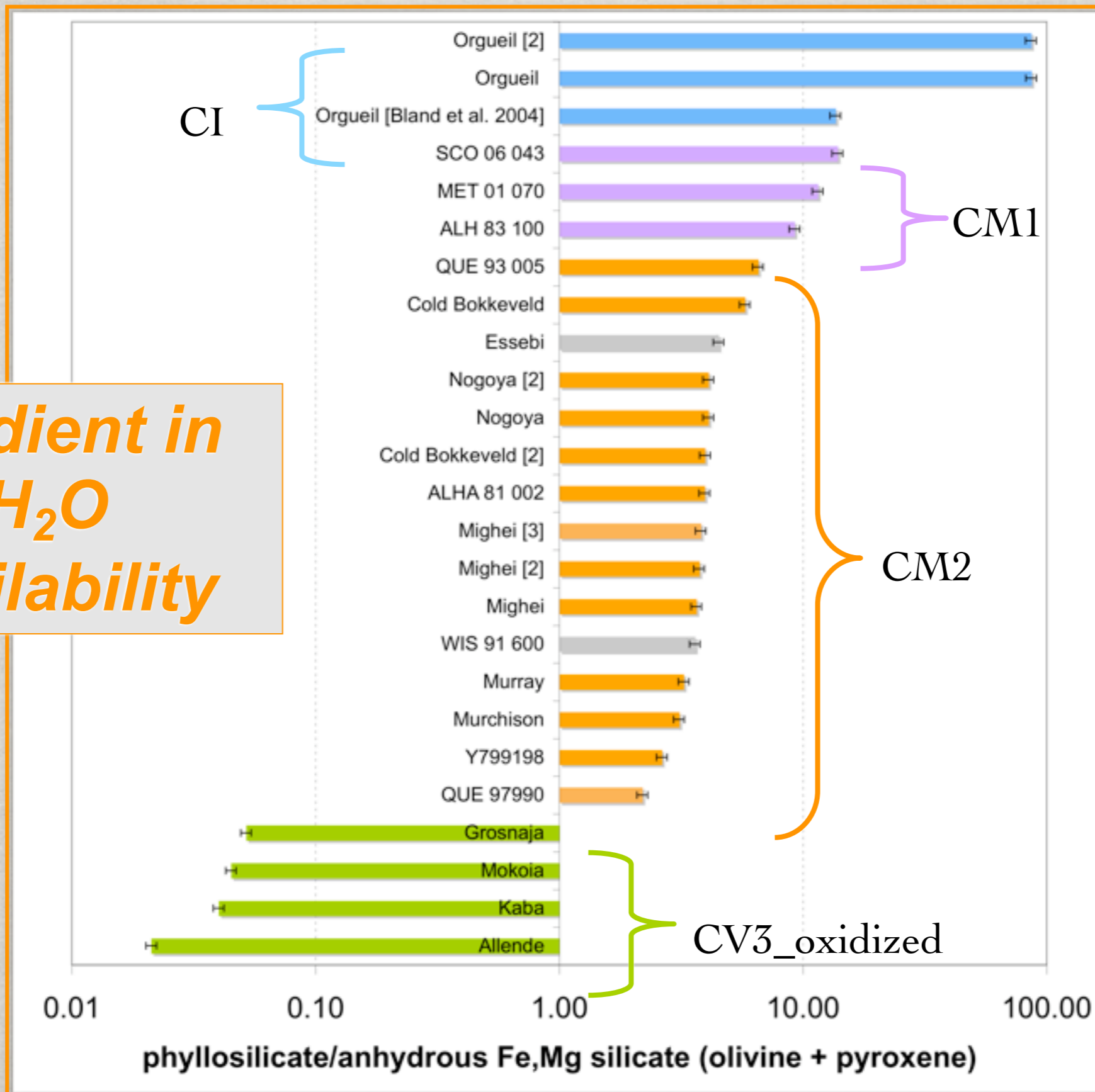
Modal Abundance



Data from Howard et al. 2009, 2010, 2011

Degree of aqueous alteration

*Gradient in
H₂O
availability*



IMPLICATIONS (?) FOR SPECTRA

- ✿ Phyllosilicates in carbonaceous chondrites (CV, CM, CI) vary in composition
 - ✿ Fe-rich generally more abundant in CV and CM2
 - ✿ CMs contain Cronstedtite - $\text{Fe}^{+2}/\text{Fe}^{+3}$ bearing
 - ✿ Mg-rich abundant in CM1 and CI
- ✿ Modal abundance varies
 - ✿ CV - < 5 vol% (but alteration could have formed anhydrous mineralogy)
 - ✿ CM2 - 65 to 80 vol%
 - ✿ CM1 - ~88 vol%
 - ✿ CI - 80 to 85 vol%